

# Introduction to CASA

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# CASA (Common Astronomy Software Applications)

- CASA is the offline data reduction package for ALMA and EVLA
- Beta Release Status:
  - in Beta Release since October 2007
  - Roughly biannual “patches”
  - register at <http://my.nrao.edu>
- Current Release: Beta Release Patch 2.4
- Next Release @Nov '09 (3.0)

# Outline

- IPython & Python
- CASA help
- CASA task interface
- MS and data selection
- The Memory Sticks
- Documentation

# CASA Interface

## • IPython

- shell access
- autoparenthesis (autocall)
- command history
- session logging
  - ipython.log – ipython command history
  - casapy.log – casa messages
- numbered input/output
- history/searching

## • Python

- Cookbook Appendix D

# Python Pointers

- to run a .py script:

- `execfile('<scriptname>')`

- example: `execfile('ngc5921_demo.py')`

- indentation matters!

- be careful when doing cut-and-paste to Python

- cut a few (4-6) lines at a time

- variables are global when using task interface!

- tasknames are objects (not variables)

# Tasks and tools in CASA

- Tasks – high-level functionality
  - function call or parameter handling interface
  - these are what you should use in tutorial
- Tools – complete functionality
  - tool.method calls, used by tasks
  - sometimes shown in tutorial scripts

# Key Tasks

- To see list of tasks after startup organized by type:

```
>tasklist
```

```
IPy:Jupyter
CASA <10>: tasklist()
Available tasks:

Import/Export      Information      Data Editing     Display/Plot
-----
importvla          listcal         flagautocorr     clearplot
(importasdm)      listhistory    flagdata         plotants
importfits        listobs        flagmanager      plotcal
importuvfits      listvis        plotxy           plotxy
exportfits        imhead         viewer
exportuvfits      imstat

Calibration        Imaging          Modelling        Utility
-----
accum             clean          setjy            help task
applycal         deconvolve    uvcontsub       help par,parameter
bandpass         feather       uvmodelfit      taskhelp
(blcal)          ft            tasklist
gaincal          invert        browsetable
fluxscale        makemask     clearplot
(fringecal)     mosaic       clearstat
clearcal
listcal
smoothcal
polcal
hanningsmooth

Image Analysis     Simulation       Single Dish
-----
imcontsub        (almasimmos)  sdaverage
imhead           sdbaseline
immath           sdcad
immoments        sdcoadd
imregrid         sdfit
imstat           sdflag
imfit            sdlist
                sdplot
                sdsave
                sdscale
                sdsMOOTH
                sdsdtat
```

# Task Execution

- two ways to invoke:
  - call from Python as functions with arguments
    - `taskname( arg1=val1, arg2=val2, ... )`
    - unspecified parameters will be defaulted (globals not used)
  - use standard tasking interface
    - use global variables for task parameters
- see Chapter 1.3 in Cookbook

# Task Interface

- standard tasking interface
  - use parameters set as global Python variables
    - `set <param> = <value>` (e.g. `vis = 'ngc5921.demo.ms'` )
  - parameter manipulation commands
    - using inp , default , saveinputs , tget
  - execute
    - `<taskname> or go` ( e.g. `clean()` )
  - return values
    - some tasks return Python dictionaries, e.g. `myval=imval()`

# Task Interface

- examine task parameters with inp :

```
IPy: Jupiter
CASA <1>: default('clean')
CASA <2>: inp('clean')
# clean :: Deconvolve an image with selected algorithm
vis                =          ''          # name of input visibility file
imagename          =          ''          # Pre-name of output images
field              =          ''          # Field Name
spw                =          ''          # Spectral windows;channels: '' is all
selectdata         =          False       # Other data selection parameters
mode               =          'mfs'       # Type of selection (mfs, channel, velocity, frequency)
niter              =          500         # Maximum number of iterations
gain               =          0.1         # Loop gain for cleaning
threshold          =          '0.0mJy'   # Flux level to stop cleaning. Must include units
psfmode            =          'clark'     # method of PSF calculation to use during minor cycles
imagermode         =          ''         # Use csclean or mosaic. If '', use psfmode
multiscale         =          []          # set deconvolution scales (pixels), default: multiscale=[] (standard CLEAN)
interactive        =          False       # use interactive clean (with GUI viewer)
mask               =          []         # cleanbox(es), mask image(s), and/or region(s) used in cleaning
imsize             =          [256, 256]  # x and y image size in pixels, symmetric for single value
cell               =          ['1.0arcsec', '1.0arcsec'] # x and y cell size, default unit arcsec
phasecenter        =          ''         # Image phase center; position or field index
restfreq           =          ''         # rest frequency to assign to image (see help)
stokes             =          'I'        # Stokes params to image (eg I,IV, QU,IQUV)
weighting          =          'natural'   # Weighting to apply to visibilities
uvtaper            =          False       # Apply additional uv tapering of visibilities.
modelimage         =          ''         # Name of model image(s) to initialize cleaning
restoringbeam      =          []         # Output Gaussian restoring beam for CLEAN image
pbcor              =          False       # Output primary beam-corrected image
minpb              =          0.1         # Minimum PB level to use
async              =          False       # If true the taskname must be started using clean(...)
CASA <3>: █
```

# Expandable Parameters

- boldface parameter are expandable

```
IPy:Jupiter
CASA <3>: tget('clean')
Restored parameters from file clean.last

CASA <4>: inp()
# clean :: Deconvolve an image with selected algorithm
vis                = 'ngc5921.usecase.ms.contsub' # name of input visibility file
imagename          = 'ngc5921.usecase.clean' # Pre-name of output images
field              = '0' # Field Name
spw                = '' # Spectral windows;channels: '' is all
selectdata       = False # Other data selection parameters
mode             = 'channel' # Type of selection (mfs, channel, velocity, frequency)
  nchan            = 46 # Number of channels (planes) in output image
  start            = 5 # first input channel to use
  width            = 1 # Number of input channels to average

niter              = 6000 # Maximum number of iterations
gain               = 0.1 # Loop gain for cleaning
threshold          = 8.0 # Flux level to stop cleaning. Must include units
psfmode           = 'clark' # method of PSF calculation to use during minor cycles
imagermode       = '' # Use csclean or mosaic. If '', use psfmode
multiscale      = [] # set deconvolution scales (pixels), default: multiscale=[] (standard CLEAN)
interactive     = False # use interactive clean (with GUI viewer)
mask               = [108, 108, 148, 148] # cleanbox(es), mask image(s), and/or region(s) used in cleaning
imsize             = [256, 256] # x and y image size in pixels, symmetric for single value
cell               = [15.0, 15.0] # x and y cell size, default unit arcsec
phasecenter        = '' # Image phase center; position or field index
restfreq           = '' # rest frequency to assign to image (see help)
stokes             = 'I' # Stokes params to image (eg I,IV, QU,IQUV)
weighting       = 'briggs' # Weighting to apply to visibilities
  robust           = 0.5 # Briggs robustness parameter
  npixels         = 0 # number of pixels to determine uv-cell size 0=> field of view

uvtaper         = False # Apply additional uv tapering of visibilities.
modelimage         = '' # Name of model image(s) to initialize cleaning
```

# Parameter Checking

- sanity checks of parameters in `inp` :

```
IPy:Jupiter
CASA <5>: psfmode='hogwarts'
CASA <6>: inp()
# clean :: Deconvolve an image with selected parameters
vis = 'ngc5921.usecase.ms.corr' # visibility file
imagename = 'ngc5921.usecase.clean' # image name
field = '0' # field name
spw = '' # spectral window
selectdata = False # select data
mode = 'channel' # mode
  nchan = 46 # Number of channels (planes) in output image
  start = 5 # first input channel to use
  width = 1 # Number of input channels to average
niter = 6000 # Maximum number of iterations
gain = 0.1 # Loop gain for cleaning
threshold = 8.0 # Flux level to stop cleaning. Must include units
psfmode = 'hogwarts' # method of PSF calculation to use during minor cycles
imagermode = '' # Use csclean or mosaic. If '', use psfmode
multiscale = [] # set deconvolution scales (pixels), default: multiscale=[] (standard CLEAN)
interactive = False # use interactive clean (with GUI viewer)
mask = [108, 108, 148, 148] # cleanbox(es), mask image(s), and/or region(s) used in cleaning
imsize = [256, 256] # x and y image size in pixels, symmetric for single value
cell = [15.0, 15.0] # x and y cell size, default unit arcsec
phasecenter = '' # Image phase center: position or field index
restfreq = '' # rest frequency to assign to image (see help)
stokes = 'I' # Stokes params to image (eg I,IV, QU,IQUV)
weighting = 'briggs' # Weighting to apply to visibilities
  robust = 0.5 # Briggs robustness parameter
  npixels = 0 # number of pixels to determine uv-cell size 0=> field of view
uvtaper = False # Apply additional uv tapering of visibilities.
modelimage = '' # Name of model image(s) to initialize cleaning
restoringbeam = [] # Output Gaussian restoring beam for CLEAN image
```

erroneous values in red

# Help on Tasks

- In-line help:

>help 'clean' OR >pdoc clean

```
IPy:Jupyter
CASA <7>: help('clean')
Help on module clean:

NAME
    clean

FILE
    /usr/lib/casapy/20.0.5444test-001/lib/python2.5/clean.py

DESCRIPTION
    # This file was generated using xslt from its XML file
    #
    # Copyright 2007, Associated Universities Inc., Washington DC
    #

FUNCTIONS
    clean_imp(vis=None, imagename=None, field=None, spw=None, selectdata=
gain=None, threshold=None, psfmode=None, imagermode=None, ftmachine=None
one, mask=None, nchan=None, start=None, width=None, imsize=None, cell=Non
er=None, outertaper=None, innertaper=None, modelimage=None, restoringbeam
r=None, cyclespeedup=None, async=None)
        Deconvolve an image with selected algorithm

        The main clean deconvolution task. It contains many functio

        1) Make 'dirty' image and 'dirty' beam (psf)
        2) Multi-frequency-continuum images or spectral channel im
        3) Full Stokes imaging
        4) Mosaicking of several pointings
        5) Multi-scale cleaning
        6) Interactive clean boxing
        7) Initial starting model

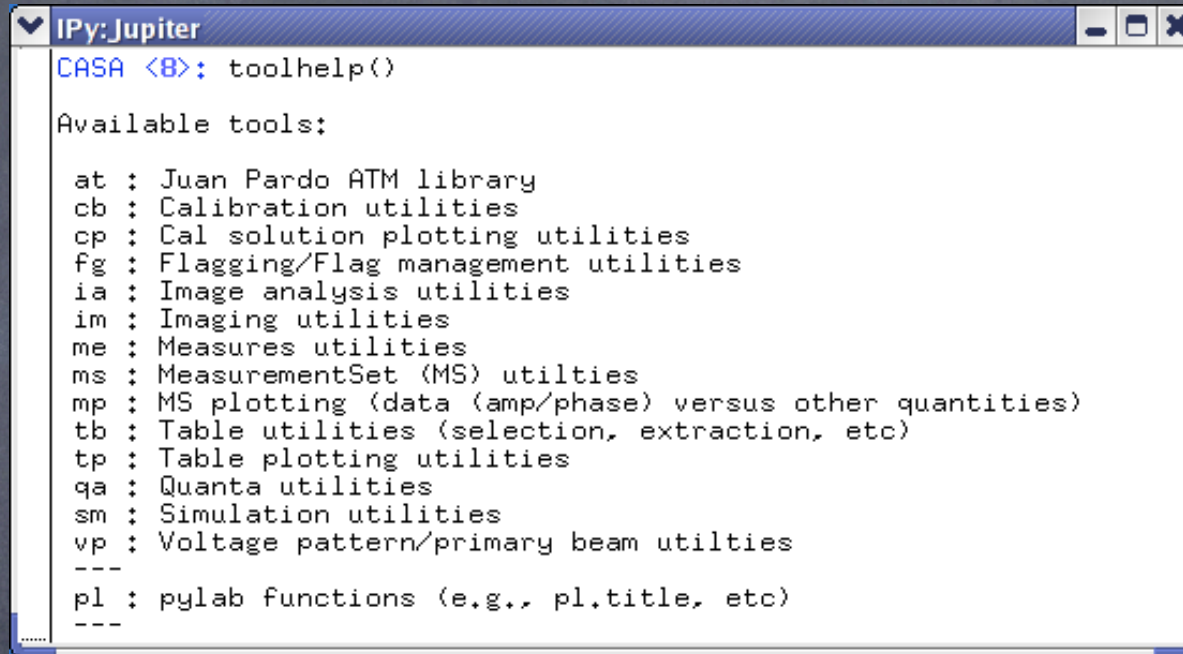
    vis -- Name of input visibility file
           default: none; example: vis='ngc5921.ms'
    imagename -- Pre-name of output images:
                 default: none; example: imagename='m2'
                 output images are:
                     m2.image: cleaned and restored image
                         With or without primary beam correction
                     m2.psf: point-spread function (dirty beam)
                     m2.flux: relative sky sensitivity over field
                     m2.model: image of clean components
                     m2.residual: image of residuals
```

# Tools in CASA

- CASA Toolkit underneath tasks
  - core AIPS++ code (mostly in C++)
- tools are functions
  - call from casapy as `<tool>.<method>()`
  - default tool objects are pre-constructed
    - e.g. imager (im) , calibrator (cb), ms (ms) , etc. (see toolhelp)

# CASA Tool List

- list of default tools from toolhelp :



```
IPy:Jupyter
CASA <8>: toolhelp()

Available tools:

at : Juan Pardo ATM library
cb : Calibration utilities
cp : Cal solution plotting utilities
fg : Flagging/Flag management utilities
ia : Image analysis utilities
im : Imaging utilities
me : Measures utilities
ms : MeasurementSet (MS) utilities
mp : MS plotting (data (amp/phase) versus other quantities)
tb : Table utilities (selection, extraction, etc)
tp : Table plotting utilities
qa : Quanta utilities
sm : Simulation utilities
vp : Voltage pattern/primary beam utilities
---
pl : pylab functions (e.g., pl.title, etc)
---
```

- tools described in the CASA User Reference Manual:

- <http://casa.nrao.edu/docs/casaref/CasaRef.html>

# The Measurement Set

- the MS is a directory on disk
  - the MAIN table in `table.*` files
  - also contains sub-tables
    - e.g. FIELD, SOURCE, ANTENNA, etc.
  - sub-tables are sub-directories
  - to copy must `cp -rf` to get contents
  - Best to remove ms with `rmtables('filename')`
    - WARNING: moving a MS can break cal-table dependencies

# Example MS

- Example: `ls`  
`ngc5921.usecase.ms`

```
smyers@olorin ~/CASA/Test $ ls ngc5921.usecase.ms
ANTENNA                POLARIZATION          table.f1              table.f3_TSM1        table.f8
DATA_DESCRIPTION       PROCESSOR             table.f10            table.f4              table.f8_TSM1
FEED                   SORTED_TABLE         table.f10_TSM1       table.f5              table.f9
FIELD                  SOURCE               table.f11            table.f5_TSM1        table.f9_TSM1
FLAG_CMD               SPECTRAL_WINDOW     table.f11_TSM1       table.f6              table.info
HISTORY                STATE                table.f2             table.f6_TSM0        table.lock
OBSERVATION           table.dat            table.f2_TSM1        table.f7
POINTING               table.f0             table.f3             table.f7_TSM1
```

- `ls ngc5921.usecase.ms/FIELD`

```
smyers@olorin ~/CASA/Test $ ls ngc5921.usecase.ms/FIELD
table.dat      table.f0      table.f0i      table.info      table.lock
```

# MAIN Table Contents

Example using task browsetable:

ngc5921.usecase.ms

	UVW	FLAG	LAG_CATEGOR	WEIGHT	SIGMA	ANTENNA1	ANTENNA2	ARRAY_ID	DATA_DESC_ID	EXPOSURE
0	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	1	1	0	0	30
1	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	27	27	0	0	30
2	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	7	7	0	0	30
3	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	2	2	0	0	30
4	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	11	11	0	0	30
5	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	17	17	0	0	30
6	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	9	9	0	0	30
7	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	19	19	0	0	30
8	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	20	20	0	0	30
9	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	18	18	0	0	30
10	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	3	3	0	0	30
11	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	15	15	0	0	30
12	[0, 0, 0]	[2, 63] Boolean	[0, 0, 0] Boolean	[23814, 23814]	[0.0514344, 0....	21	21	0	0	30

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Browsing table: ngc5921.usecase.ms

# Data Selection Example

- standard selection parameters

- e.g. for task gaincal:

```
CASA <14>: inp
-----> inp()
# gaincal :: Determine temporal gains from calibrator observations:

vis                = 'ngc5921.ms'      # Name of input visibility file
caltable           = 'ngc5921.gcal'    # Name of output calibration table
field              = '0,1'            # field names or index of calibrators ''=>all
spw                = '0:2~56'        # spectral window:channels: ''=>all
selectdata       = True              # Other data selection parameters
  timerange        = ''               # time range: ''=>all
  uvrange          = ''               # uv range''=all
  antenna          = ''               # antenna/baselines: ''=>all
  scan             = ''               # scan numbers
  msselect         = ''               # Optional data selection (Specialized. but see help)
```

- field and spw common standard selections
    - expandable selectdata with other selections as sub-parameters

# Data Selection Syntax

- see Chapter 2.5 of Cookbook
  - field – string with source name or field ID
    - can use '\*' as wildcard, first checks for name, then ID
    - example: field = '1331+305' ; field = '3C\*' ; field = '0,1,4~5'
  - spw – string with specwindow ID plus channels
    - use ':' as separator of spw from optional channelization
    - use '^' as separator of channels from step/width
    - example: spw = '0~2' ; spw = '1:10~30' ; spw = '2~5:5~54^5'

# Selection Syntax

- see Chapter 2.5 of Cookbook
  - antenna - string with antenna name or ID
    - first check for name, then ID (beware VLA name 1-27, ID 0-26)
    - example: antenna = '1~5,11' ; antenna = 'VA\*'
  - timerange - string with date/time range
    - specify 'T0~T1' , missing parts of T1 default to T0, can give 'T0+dT'
    - example: timerange = '2007/10/16/01:00:00~06:30:00'

# The Memory Sticks etc.

## • Contents:

- "Install\_README" - installation instructions
- "installation" - directory contains 32-bit linux and intel mac files. Please see me if you need 64-bit linux. Please install before Wed.
- "casa\_scripts\_and\_data.tar" copy this to your hard drive before Wed. Note a copy of the casa cookbook is included in this directory. A slightly newer version is available at [http://www.aoc.nrao.edu/~smyers/naug/doc/cookbook/casa\\_cookbook.pdf](http://www.aoc.nrao.edu/~smyers/naug/doc/cookbook/casa_cookbook.pdf)

## • Additional CASA Scripts & Data Page

- <http://casa.nrao.edu/casatraining.shtml>

# Getting User Support

- First stop:

- CASA Home: <http://casa.nrao.edu>
- Cookbook, and "help" within CASA

- CASA Helpdesk & Support

- "Helpdesk" at <http://my.nrao.edu> (will migrate to more user friendly helpdesk in future)
- You will need to register first
  - Submit questions, suggestions, bugs

# CASA Documentation

- CASA Analysis cookbook:

- [http://casa.nrao.edu/Doc/Cookbook/casa\\_cookbook.pdf](http://casa.nrao.edu/Doc/Cookbook/casa_cookbook.pdf)

- CASA User Reference Manual:

- <http://casa.nrao.edu/docs/casaref/CasaRef.html>

- Python:

- <http://python.org/doc> (e.g., see Tutorial for novices)

- IPython:

- <http://ipython.scipy.org/moin/Documentation>

- matplotlib:

- <http://matplotlib.sourceforge.net/>